



LOOK BEFORE YOU LEAP

Using the Solar Gravitational Lens to explore exoplanets

The scale of the problem



Extrasolar scales

- Distance to the nearest star is almost 300,000 AU
- That is almost 100 million times the famous “one giant leap”
- If the Earth is a grape and the Moon is a peppercorn a foot away...
- ... the nearest star would be more than halfway around the Earth.
- It makes sense to look before we take such a giant leap

What does it entail to look?

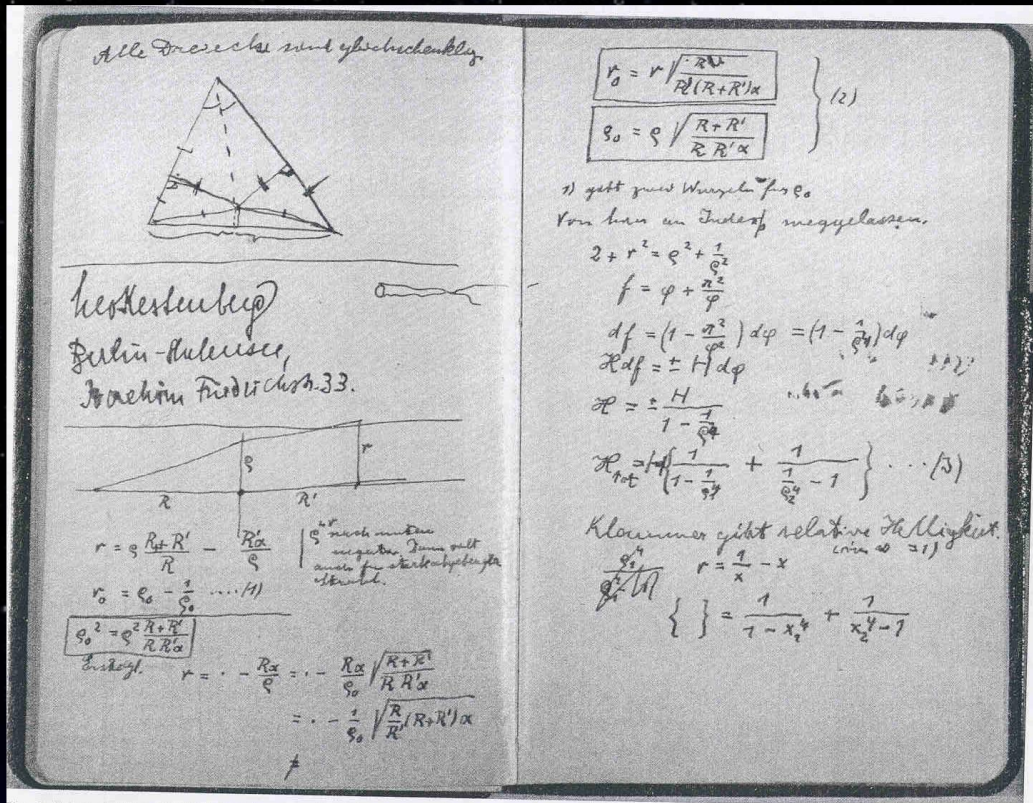
- Fundamentally, three issues:
 - Resolution: We want to be able to see details of a distant world. Diffraction
 - Brightness: We want enough light to form an image. Photons are scarce
 - Noise: There are many contaminating sources of light. Noise can get amplified

The tyranny of the diffraction limit

- Angular resolution is roughly proportional to λ/d
- Resolving features of 10 km at 10 light years implies $d \sim 10,000$ km
- Large baselines and interferometry may help but there's noise
 - Light contamination from host star
 - Exozodiacal noise
- Other techniques (e.g., rotational deconvolution) have limitations

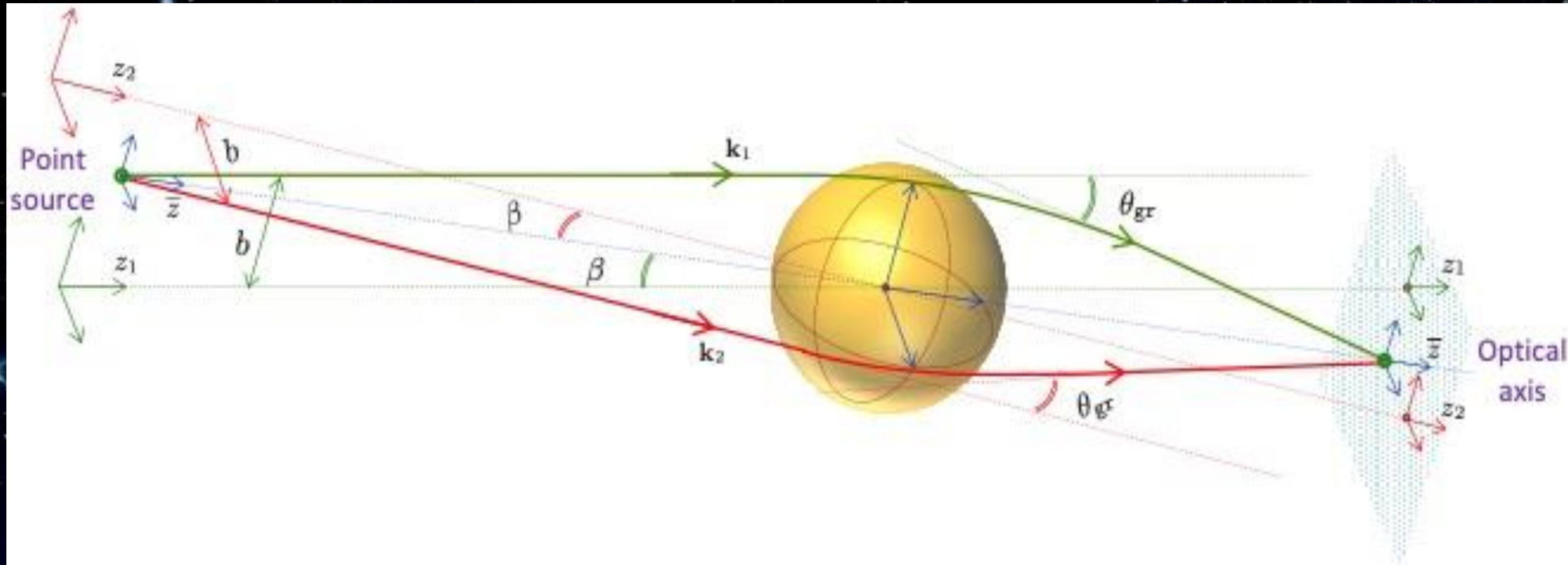
The Sun to the rescue!

- The Sun bends rays of light and acts as an imperfect lens



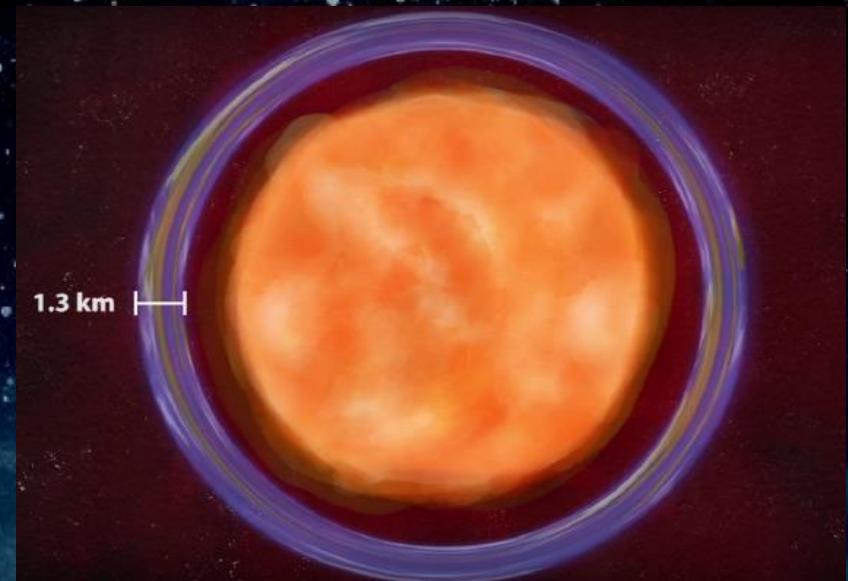
The focal region

- Parallel rays of light are deflected, meeting at 550+ AU



Projection vs. view

- The Sun projects a km-scale blurry image of a distant object
- A telescope in the focal region only sees a thin Einstein-ring



Images must be reconstructed

- Multiple observations needed
- Each observation captures 1 “pixel”
- What is imaged is a moving target
- Long exposure times are required
- Precise stationkeeping, location tracking essential

Why not go there today?

- Target must be known in advance to pick focal region
- Imaging has many challenges
- Distance is enormous

The imaging challenge: signal-to-noise

- From any one vantage point in the image plane, all we see is a very faint Einstein ring through the very bright solar corona
- Corona is $10^4 - 10^5$ times brighter than the Einstein ring
- Even if the corona background is removed, we have to deal with shot noise due to the very low signal photon count
- We must also block light from the Sun
- There will be light from the host star, brightening the Einstein ring in specific places
- Light from interlopers must be known

The imaging challenge: collecting light

- Low SNR and stochastic (shot) noise has one solution: increased integration times
- Prolongs mission duration
- Represents added navigational challenge
- Must deal with a temporally changing target

The imaging challenge: imperfect lens

- The monopole SGL is a lens with significant spherical aberration
- Quadrupole contributes significant astigmatism
- Reconstruction (deconvolution) is possible but it substantially amplifies noise
- Without deconvolution, a less noisy but blurry image results
- Trade-off between noise, resolution and sharpness must be considered

The imaging challenge: moving target

- An Earth-like planet rotates, changes significantly over the scale of minutes
- Its illumination changes with the seasons
- Its appearance changes: short-term stochastic changes (weather) and periodic changes (e.g., ice cover, vegetation)

Recovering a moving target

The RIBFED experiment -- temporal testing ground

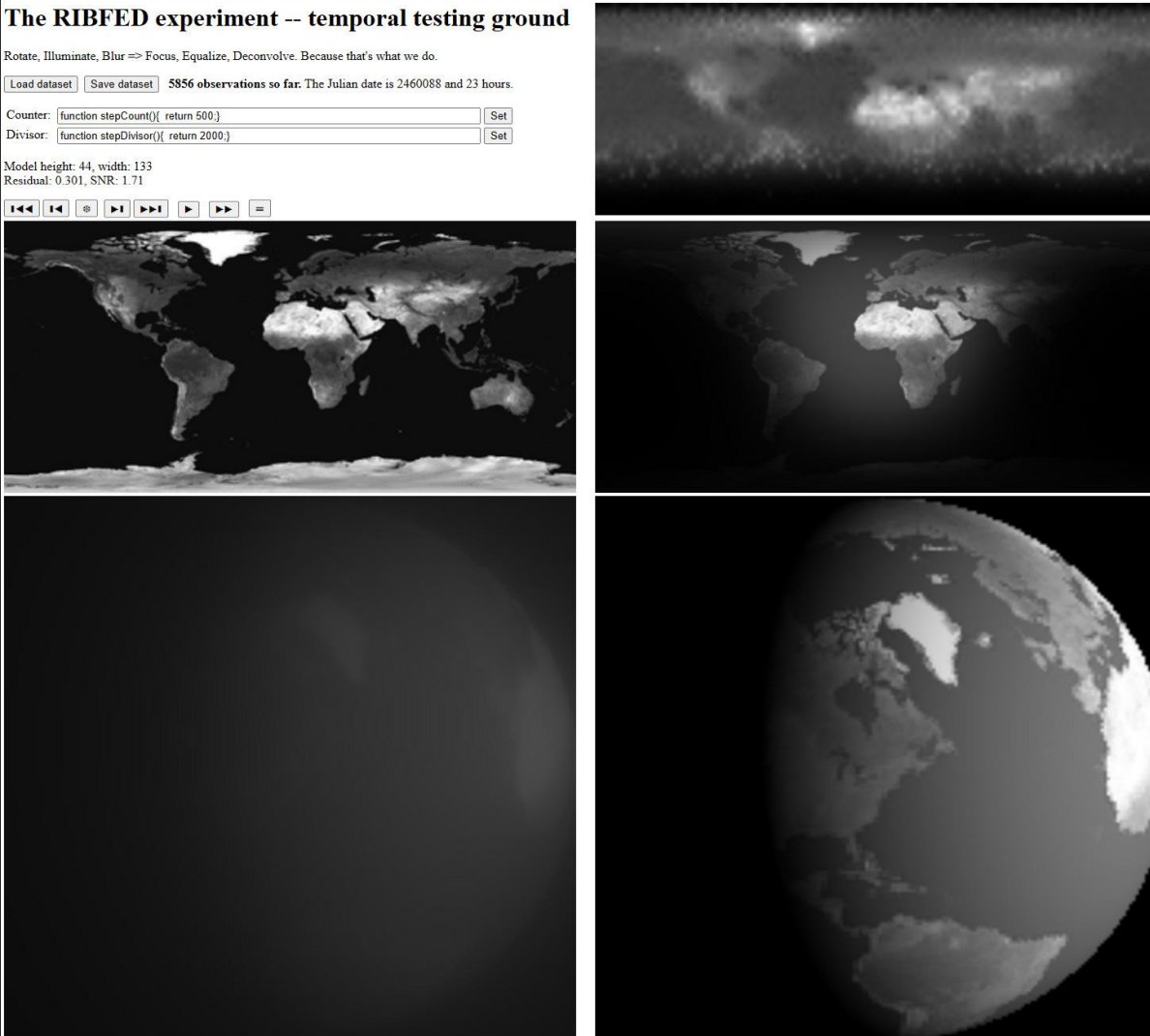
Rotate, Illuminate, Blur \Rightarrow Focus, Equalize, Deconvolve. Because that's what we do.

5856 observations so far. The Julian date is 2460088 and 23 hours.

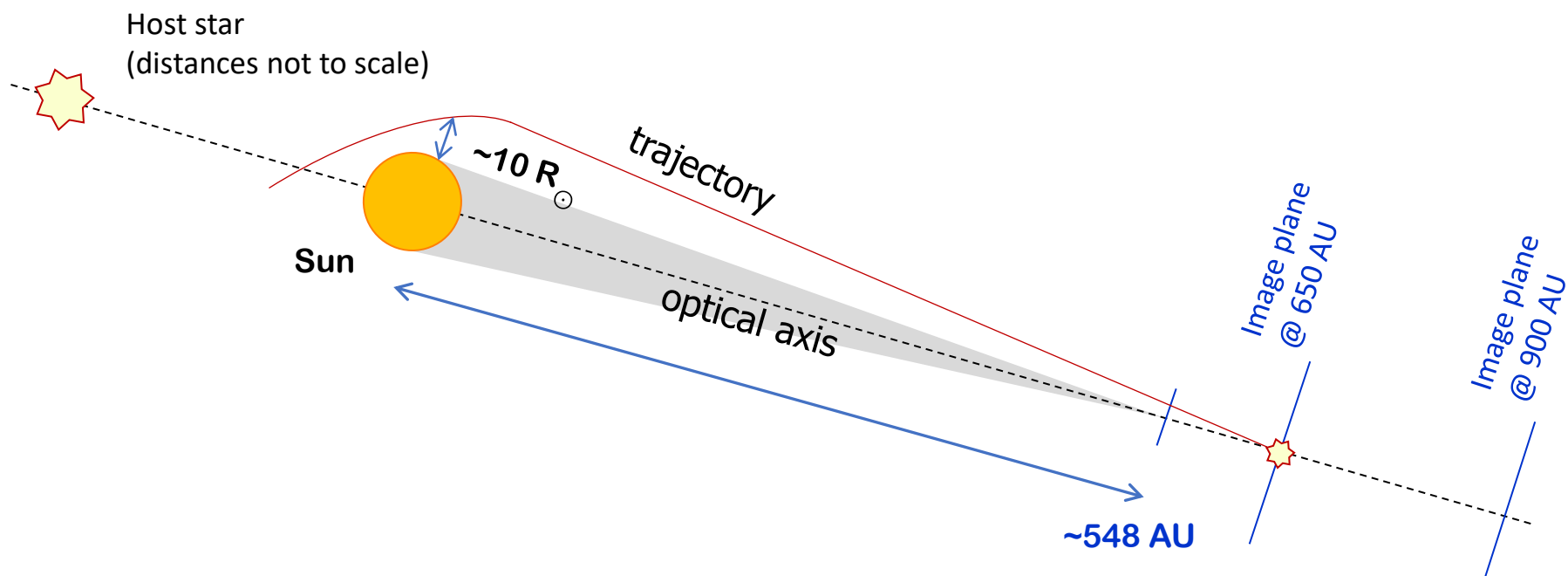
Counter:

Divisor:

Model height: 44, width: 133
Residual: 0.301, SNR: 1.71

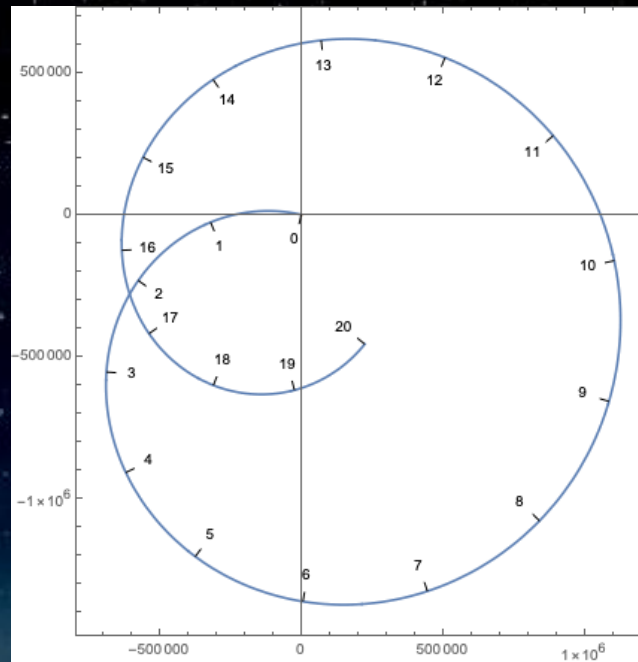


How to get there?



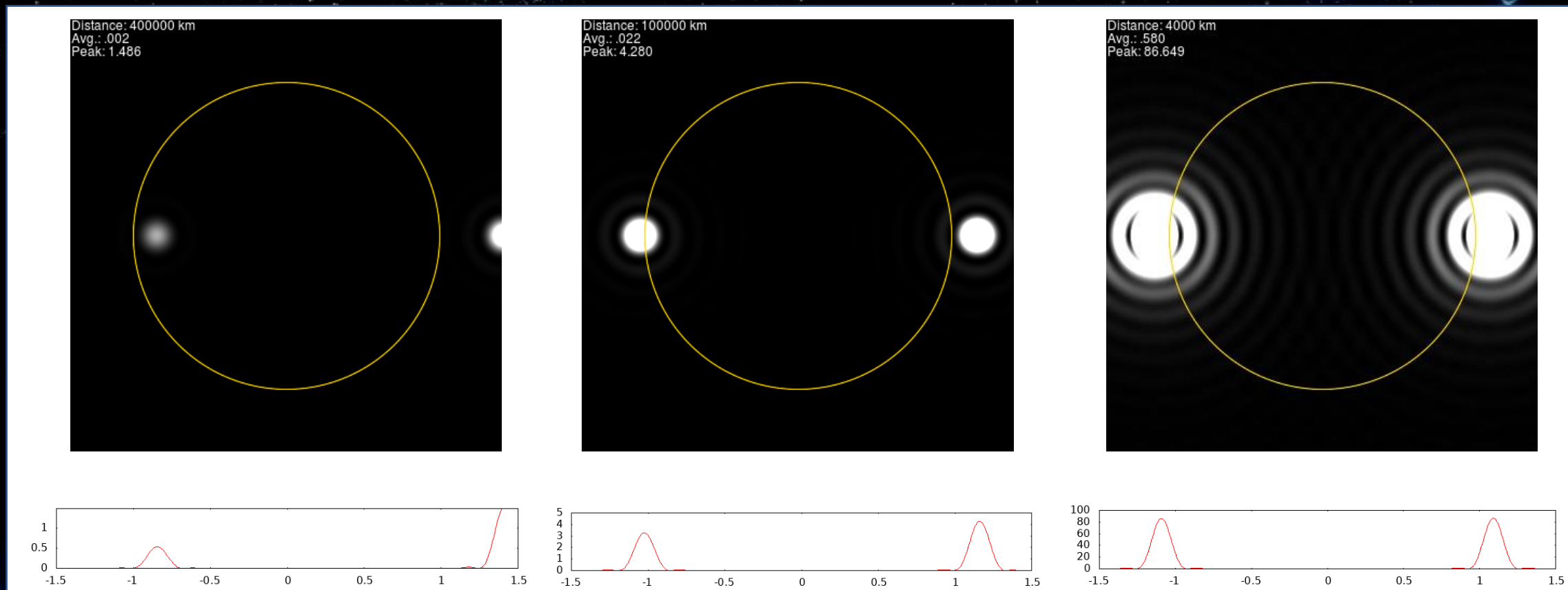
Navigation is a serious challenge

- Host star is visible throughout the cruise
- Focal region wobbles by as much as a million km



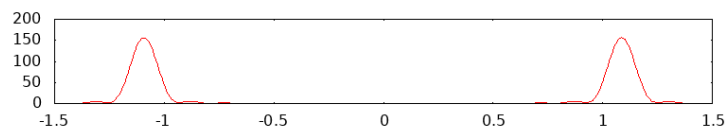
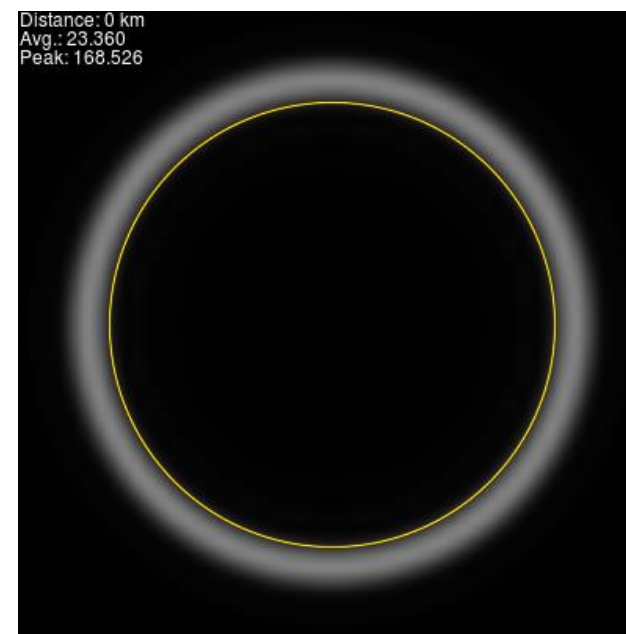
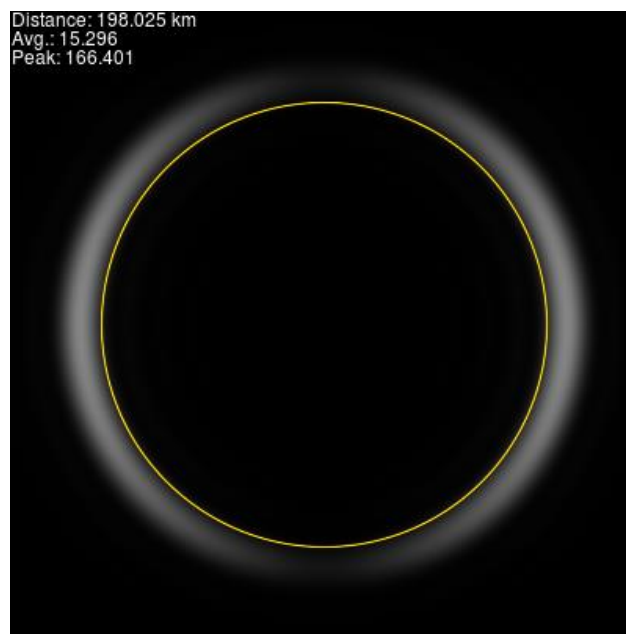
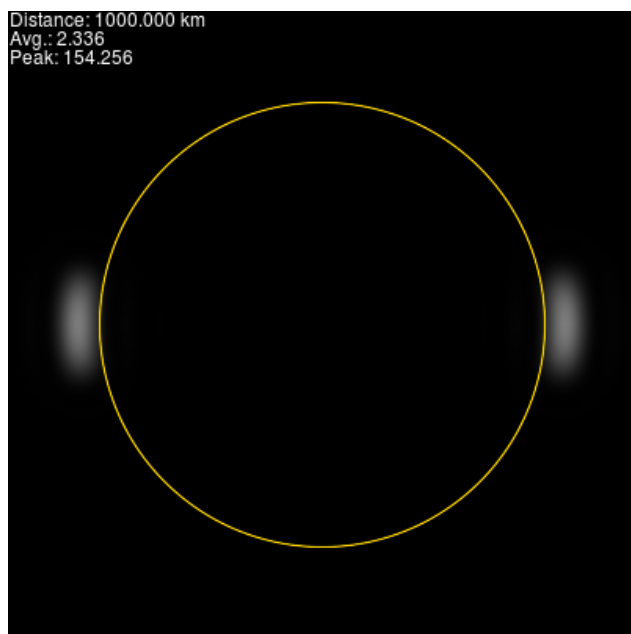
Navigating from 400,000 to 4,000 km

- Light amplification by $O(100)$



Zeroing in on the host star

- From 1,000 km to the exostar optical axis



Fine-grained navigation

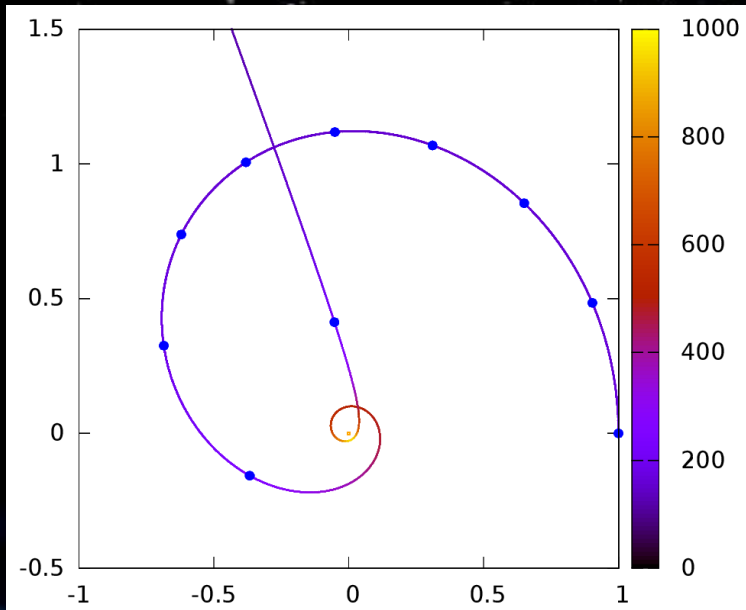
- Exoplanet image located $O(10,000 \text{ km})$ from host star image in the image plane
- Approximate direction known
- Approaching the exoplanet image is a repeat of the host star approach but on a smaller and fainter scale, on the host star light background
- Exoplanet image size is of $O(1 \text{ km})$

Need for a local reference frame

- Projected image motion is known but it is noninertial
- Absolute navigation at that accuracy is unrealistic
- Once the exoplanet image is located, two or more telescopes could establish its precise boundary
- These telescopes establish and maintain a local reference frame
- Observing telescope can move from pixel to pixel on a meter-scale grid with respect to this frame
- Primary observable: Total amount of light collected from the Einstein-ring at each pixel location

The means to get there

- Sundivers!



Technical challenges

- Large spacecraft vs. in-flight assembly
- Systems to survive 25-30 years cruise plus up to 10 years science
- Autonomous constellation (required for navigation, corona light suppression)
- Coronagraph (possible use of discarded sails as external sunshades?)
- Power (40+ years)
- Communication (650 – 900 AU)

Thank you

- Questions?

